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GPO PRICE \$ \_\_\_\_\_\_

CFSTI PRICE(S) \$ \_\_\_\_\_

Hard copy (HC) \_\_\_\_\_\$ 1.00

Microfiche (MF) \_\_\_\_\_\_ 150

ff 653 July 65

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TE FOR	39	(THRU)
A GIL	(NASA CR OR TMX OR AD NUMBER)	(CODE)

# THE TIME-HISTORY OF A SATELLITE AROUND AN OBLATE PLANET

**FEBRUARY 1965** 

M.C. ECKSTEIN
SCIENTIST SPECIALIST
ADVANCE FLIGHT MECHANICS DEPARTMENT

Y.Y. SHI

SCIENTIST SPECIALIST
ADVANCE FLIGHT MECHANICS DEPARTMENT

J. KEVORKIAN

CONSULTANT

DOUGLAS AIRCRAFT COMPANY, INC.



DOUGLAS MISSILE & SPACE SYSTEMS DIVISION

## THE TIME-HISTORY OF A SATELLITE

## AROUND AN OBLATE PLANET

M. C. Eckstein, Y. Y. Shi to and J. Kevorkian

## ABSTRACT

In the previous work on the motion of a close satellite around an oblate planet, the orbital elements and perturbations have been described as functions of the central angle between the instantaneous node and the satellite. A complete theory, however, requires the elements and perturbations as functions of the time. Therefore the relationship between time and the central angle between node and satellite is presented here. Although the problem is mathematically nothing but a quadrature, it is practically quite complicated because the evaluation of the occurring integral requires some rather lengthy algebraic manipulations. Part of the calculations are avoided by using a new technique to numerically evaluate certain coefficients that depend only on the initial conditions. Furthermore, the energy integral is used to evaluate certain terms that otherwise would require consideration of higher order terms in the differential equations.

## CREDIT

Prepared under NASA Contract No. NASw-901

<sup>†</sup>Scientist Specialist, Advance Flight Mechanics Department

<sup>††</sup> Scientist Specialist, Advance Flight Mechanics Department; Assistant Professor of Aeronautics, Carnegie Institute of Technology, Pittsburgh, Pennsylvania.

Consultant to Douglas Aircraft Company, Inc.; Assistant Professor of Aeronautics and Astronautics, University of Washington, Seattle, Washington.

## 1. INTRODUCTION

Eckstein, Shi, and Kevorkian (1964) discussed the geometry of the orbit of a satellite around an oblate planet. Particular emphasis was given to the nature of the orbit near "the critical inclination" and to the question of defining the motion for all inclinations by means of uniform asymptotic expansions. This paper is a continuation of the mentioned reference and discusses the time-history of the motion. The notation and definitions of the cited reference will be used throughout, and equations shown there will be referred to by the corresponding number followed by an asterisk.

The solution of the time-history of the motion is formally very simple, since (cf. (3.5\*)) it reduces to a quadrature once the orbit has been defined. However, the practical problem of exhibiting an explicit analytic formula for the time, correct to some order, is quite involved for the following reasons.

Inclusion of short-period terms in the solution introduces considerable algebraic complexity in this problem. Struble (1962) neglected these terms by using an averaged differential equation as a starting point for calculating the first approximation to the long-period and secular terms in the time-history. In this paper all terms in the solution are retained in order to define the time-history completely to order  $\varepsilon$ . The presence of perturbations of course eliminates the advantage of using the eccentric anomaly as an independent variable. Therefore, short-period terms are expressed by their convergent Fourier series expansions in terms of  $\phi$ , the central angle between the instantaneous node and the radius vector.

In addition to the algebraic complexity of the problem where short-period terms are included, the inherent difficulty associated with the quadrature of the equation for  $dt/d\phi$  is the presence of long-period terms on the right hand side. (These are trigonometric terms whose arguments are integer multiples of  $\epsilon\phi$ .) Since such terms drop in order by one power of  $\epsilon$  after integration, it is necessary to consider terms of order  $\epsilon^{n+1}$  in order to include all long-period terms of order  $\epsilon^n$  in the solution. Thus, the major difficulty in deriving the solution for t correct to order  $\epsilon$  stems from the presence of certain unknown long-period terms of order  $\epsilon^2$  in  $dt/d\phi$ .

To evaluate these terms would in general require consideration of perturbations of order  $\varepsilon^3$  for the inclination and reciprocal radius. However, in this particular case it is shown that the existence of the energy integral is sufficient to define all long-period terms that arise to order  $\varepsilon^2$ .

In order to avoid excessive algebraic manipulations, some of the coefficients in the present solution are not evaluated literally. These coefficients occur in expressions where the general form is known and they may be expressed either as definite integrals or as the solutions of systems of linear algebraic equations. As these coefficients are constant for any given motion, one need only evaluate them numerically once for each set of initial conditions.

Briefly, this is due to the fact that the long-period behavior of a term of order  $\epsilon^n$  is governed by boundedness criteria on terms of order  $\epsilon^{n+1}$ .

## 2. DISCUSSION

# 2.1 The Differential Equation for the Time

Struble (1960) gives the following general result for the rate of change of \$\phi\$ for a problem with an arbitrary cylindrically symmetric potential.

$$\frac{d\phi}{dt} = \frac{pu^2}{\cos i} + \frac{\cos^3 i \cos \theta}{p \sin^2 i \sin \theta} \frac{\partial U}{\partial \theta}$$
 (2.1)

After substituting for  $\theta$  and  $\frac{\partial U}{\partial \theta}$  from (3.5h\*) and (3.2\*) and some manipulation one obtains

$$\frac{dt}{d\phi} = \frac{\cos i}{pu^2} - \varepsilon \left[ 2 \frac{\cos^5 i}{p^3 u} \sin^2 \phi \right] + \varepsilon^2 \left[ 4 \frac{\cos^9 i}{p^5} \sin^4 \phi + 28 \frac{cu}{p^3} \cos^5 i \sin^2 i \sin^4 \phi \right]$$

$$- 12 \frac{cu}{p^3} \cos^5 i \sin^2 \phi \right] + O(\varepsilon^3)$$
(2.2)

Upon substitution of the expansions for i and u into the right hand side, (2.2) reduces to

$$\frac{dt}{d\phi} = \frac{\cos i_0}{pu_0^2} - \varepsilon \left[ \frac{\sin i_0}{pu_0^2} i_1 + 2 \frac{\cos i_0}{pu_0^3} u_1 + 2 \frac{\cos^5 i_0}{p^3 u_0} \sin^2 \phi \right]$$

$$+ \varepsilon^2 \left[ -\frac{1}{2} \frac{\cos i_0}{pu_0^2} i_1^2 - \frac{\sin i_0}{pu_0^2} i_2 + 2 \frac{\sin i_0}{pu_0^3} i_1^2 u_1 + 3 \frac{\cos i_0}{pu_0^4} u_1^2 \right]$$

$$- 2 \frac{\cos i_0}{pu_0^3} u_2 + 10 i_1 \frac{\cos^4 i_0 \sin i_0}{p^3 u_0} \sin^2 \phi + 2 \frac{\cos^5 i_0}{p^3 u_0^2} u_1 \sin^2 \phi$$

$$+ 4 \frac{\cos^{9} i_{o}}{p^{5}} \sin^{4} \phi + 28 \frac{cu_{o}}{p^{3}} \cos^{5} i_{o} \sin^{2} i_{o} \sin^{4} \phi$$

$$- 12 \frac{cu_{o}}{p^{3}} \cos^{5} i_{o} \sin^{2} \phi] + 0(\epsilon^{3})$$
(2.3)

All terms of the order unity and  $\varepsilon$  on the right hand side of (2.3) are given in the original reference of the authors. However, as mentioned previously, a solution for t to order  $\varepsilon$  must also include consideration of the long-period terms proportional to  $\varepsilon^2$  in (2.3). This necessitates the evaluation of  $i_2$  and  $u_2$  occurring to order  $\varepsilon^2$  in (2.3). The lengthy expressions for these two quantities which had not been computed earlier are shown in Appendix 1. In addition, the leading term of (2.3), viz.  $\cos i_0 / \operatorname{pu}_0^2$ , contains the unknown terms  $i_{02}$  and  $e_2$  to order  $\varepsilon^2$ . As was pointed out in the original reference the determination of these terms requires the knowledge of the differential equations for i and e correct to order  $\varepsilon^3$ . It will be shown in the next section that for the present case, one can deduce the form in which these terms occur by using the energy integral.

# 2.2 Use of the Energy Integral to Compute Long-Period Terms

The energy integral

$$\frac{1}{2} \left( \frac{d\mathbf{r}}{dt} \right)^2 + \frac{1}{2} r^2 \left( \frac{d\psi}{dt} \right)^2 \sin^2 \theta + \frac{1}{2} r^2 \left( \frac{d\theta}{dt} \right)^2 - U = E$$
 (2.4)

can be brought to the following exact form in terms of the present variables

$$u^{\frac{1}{4}} \left(\frac{dt}{d\phi}\right)^{2} = \left(\frac{du}{d\phi}\right)^{2} \left[2E + u\left(2 - \frac{p^{2}}{\cos^{2}i}u\right) + \frac{2}{3}\varepsilon u^{3} \left(1 - 3\sin^{2}i\sin^{2}\phi\right) + \varepsilon^{2}cu^{5} \left(1\frac{1}{4}\sin^{4}i\sin^{4}\phi - 12\sin^{2}i\sin^{2}\phi + \frac{6}{5}\right]^{-1}$$
(2.5)

Here, E is the energy constant expressible in terms of the initial conditions, and the potential U, given by  $(3.2^{\circ})$ , has been used. It is clear that knowledge of the orbit (i.e.,  $u(\phi, \epsilon)$ ,  $i(\phi, \epsilon)$ ) together with  $dt/d\phi$ , as given by (2.3) to any order  $\epsilon^n$  should lead to an identity to  $O(\epsilon^n)$  when these values are substituted into (2.5). Thus, one could use (2.5) to check the solution to any order. In the present case however, a converse use of the energy integral will be invoked. The orbit defined by u and i is known to  $O(\epsilon^2)$  and it will be assumed that these results are free of algebraic errors. The energy integral will then be used to define those terms of  $dt/d\phi$  which are unknown to the order the calculations have been carried out. A partial check of the accuracy of the results used for u and i will be the cancellation of all known terms after substitution into (2.5) leaving only a definition of the unknown long-period terms one seeks.

As was pointed out earlier, the long period terms of order  $\epsilon^2$  in the leading term of (2.3) are unknown. To evaluate these,  $\cos i_0$  /  $pu_0^2$  is expanded in its Fourier series as shown below

$$\frac{\cos i_0}{pu_0^2} = \frac{p^3}{\cos^3 i_0} \left[ \frac{1}{(1-e^2)^{3/2}} + \frac{2}{(1+e^2)^{3/2}} \sum_{k=1}^{\infty} \left( \frac{\sqrt{1-e^2}-1}{e} \right)^k \cdot (1+k\sqrt{1-e^2}) \cos k \left( \phi - \omega \right) \right]$$
 (2.6)

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The terms under the summation sign in the above have a short period and hence need not be evaluated to  $0(\varepsilon^2)$ . Conversely, it is necessary to compute the form of the first term,  $p^3/\cos^3i_o$   $(1-e^2)^{3/2}$  correct to  $0(\varepsilon^2)$  since this term depends only on  $\tilde{\phi}$  (or  $\tilde{\phi}$ ). It will be shown in what follows that substitution of all known terms into the energy integral leads to precisely the expression for the above term.

When  $dt/d\phi$  is eliminated from (2.5) by using (2.2) one obtains

After substituting the expansions for the variables on the right hand side of (2.7) in the form assumed in the original reference one obtains after some manipulation and ordering

$$2 \frac{\cos^{2}i_{o}}{p^{2}} E + u_{o} \left(2 \frac{\cos^{2}i_{o}}{p^{2}} - u_{o}\right) - \left(\frac{\partial u_{o}}{\partial \phi}\right)^{2} + \varepsilon \left\{-2 \frac{\partial u_{1}}{\partial \phi} \frac{\partial u_{o}}{\partial \phi} - 2 \frac{\partial u_{o}}{\partial \phi} \frac{\partial u_{o}}{\partial \phi} - 2 \frac{\partial u_{o}}{\partial \phi} \frac{\partial u_{o}}{\partial \phi}\right\}$$
$$-2 \frac{\sin^{2}i_{o}}{p^{2}} i_{1} \left(E + u_{o}\right) + 2u_{1} \left(\frac{\cos^{2}i_{o}}{p^{2}} - u_{o}\right) - \frac{1}{4} \frac{\cos^{6}i_{o}}{p^{4}} u_{o} \left[2E\right]$$
$$+ u_{o} \left(2 - \frac{p^{2}}{\cos^{2}i_{o}} u_{o}\right) \left[\sin^{2}\phi + \frac{2}{3} \frac{\cos^{2}i_{o}}{p^{2}} u_{o}^{3} \left(1 - 3 \sin^{2}i_{o} \sin^{2}\phi\right)\right]$$

$$+ \varepsilon^{2} \left\{ -\left(\frac{\partial u_{1}}{\partial \phi}\right)^{2} - 2 \frac{\partial u_{1}}{\partial \phi} \frac{\partial u_{0}}{\partial \phi} - \left(\frac{\partial u_{0}}{\partial \phi}\right)^{2} - 2 \frac{\partial u_{2}}{\partial \phi} \frac{\partial u_{0}}{\partial \phi} - 2 \frac{\partial u_{1}}{\partial \phi} \frac{\partial u_{0}}{\partial \phi} \right\}$$

$$- 2 \frac{\sin 2i_{0}}{p^{2}} (E + u_{0})i_{2} - 2 \frac{\cos 2i_{0}}{p^{2}} (E + u_{0})i_{1}^{2} - 2 \frac{\sin 2i_{0}}{p^{2}} u_{1}i_{1}$$

$$+ 2 \left(\frac{\cos^{2}i_{0}}{p^{2}} - u_{0}\right)u_{2} - u_{1}^{2} + 48 \frac{\cos^{5}i_{0} \sin i_{0}}{p^{4}} i_{1}u_{0}(E + u_{0}) \sin^{2}\phi \right)$$

$$- 8 \frac{\cos^{6}i_{0}}{p^{4}} (E + 2u_{0})u_{1} \sin^{2}\phi - 16 \frac{\cos^{3}i_{0} \sin i_{0}}{p^{2}} i_{1}u_{0}^{3} \sin^{2}\phi$$

$$+ 12 \frac{\cos^{4}i_{0}}{p^{2}} u_{0}^{2}u_{1} \sin^{2}\phi - \frac{2}{3} \frac{\sin 2i_{0}}{p^{2}} i_{1}u_{0}^{3} - \frac{\sin 4i_{0}}{p^{2}} i_{1}u_{0}^{3} \sin^{2}\phi$$

$$+ 2 \frac{\cos^{2}i_{0}}{p^{2}} u_{0}^{2}u_{1} (1 - 3 \sin^{2}i_{0} \sin^{2}\phi)$$

$$+ \left[2E + u_{0}\left(2 - \frac{p^{2}}{\cos^{2}i_{0}} u_{0}\right)\right]\left[12 \frac{\cos^{10}i_{0}}{p^{6}} u_{0}^{2} \sin^{4}\phi \right]$$

$$+ 56c \frac{\cos^{6}i_{0}}{p^{4}} u_{0}^{3} \sin^{2}i_{0} \sin^{4}\phi - 24c \frac{\cos^{6}i_{0}}{p^{4}} u_{0}^{3} \sin^{2}\phi \right]$$

$$- \frac{8}{3} \frac{\cos^{6}i_{0}}{p^{4}} u_{0}^{4} \sin^{2}\phi (1 - 3 \sin^{2}i_{0} \sin^{2}\phi + 14 \sin^{4}i_{0} \sin^{4}\phi) \right\} = 0(\varepsilon^{3})$$

$$+ cu_{0}^{5} \frac{\cos^{2}i_{0}}{p^{2}} \left(\frac{6}{5} - 12 \sin^{2}i_{0} \sin^{2}\phi + 14 \sin^{4}i_{0} \sin^{4}\phi \right) \right\} = 0(\varepsilon^{3})$$

In the above the known expressions for u<sub>o</sub>, u<sub>1</sub>, u<sub>2</sub>, i<sub>1</sub>, and i<sub>2</sub> have not yet been used. If these results are now substituted into (2.8), all short-period terms cancel identically as expected, and what remains is the following expression for the long-period term under consideration.

$$\frac{\cos^{2} i_{o}}{p^{2}} (1 - e^{2}) = -2E + \epsilon \frac{\cos^{6} i_{o}}{p^{6}} [(\frac{1}{3} + e^{2})(1 - 3\cos^{2} i_{o}) + \frac{e^{2}}{4} (1 + \cos^{2} i_{o})\cos 2\omega] + \epsilon^{2} [\eta_{o} + \eta_{o}\cos 2\omega + \eta_{h}\cos 4\omega] + 0(\epsilon^{3})$$
(2.9)

Actually, what is needed is the negative three-halves power of (2.9) correct to  $O(\epsilon^2)$ . This can be easily computed once the quantities denoted by  $\eta_0$ ,  $\eta_2$ , and  $\eta_h$  have been defined.

Before discussing the evaluation of  $n_0$ ,  $n_2$ , and  $n_4$ , it is pointed out that in the term of order  $\varepsilon$  in (2.9), e and  $i_0$  have not been expanded because the higher order terms in these expansions are no longer negligible. In contrast, for the evaluation of (3.15\*) and (3.20\*) e and  $i_0$  or (e\* and  $i_0$ \*) were replaced by  $e_0$ \* and  $i_0$ \* since the solution was valid only to  $O(\varepsilon)$  there. However, one need only use  $e_0$  and  $i_0$ 0 for the values of e and  $i_0$ 0 occurring in the term proportional to  $\varepsilon^2$  in (2.9).

Thus  $\eta_0$ ,  $\eta_1$ , and  $\eta_4$  formally depend on the initial values  $e_0$  and  $i_{00}$ . The form of the term of  $O(\epsilon^2)$  in (2.9) can be deduced by inspection of the terms of  $O(\epsilon^2)$  in (2.8). Each of these may be expressed as the following double sum

$$\sum_{n=0}^{8} \sum_{k=-2}^{2} a_{nk} \cos[n(\phi - \omega) + 2k\omega]$$

where the a are functions of e and i oo

Now the identity (2.8) must be independent of  $\phi$  to any order in  $\epsilon$ . In particular, to  $O(\epsilon^2)$  all the contributions  $a_{nk} \cos[n(\phi - \omega) + 2k\omega]$  must cancel when  $n \neq 0$  leaving only terms which can be brought to the form shown in (2.9).

Unfortunately, the explicit evaluation of  $\eta_0$ ,  $\eta_2$ , and  $\eta_4$  in terms of  $i_{00}$  and  $e_0$  requires extremely involved algebraic calculations. Furthermore, since the energy integral has already been used in arriving to this stage there is no independent procedure for checking the results if one were to actually derive explicit formulae for  $\eta_0$ ,  $\eta_2$ , and  $\eta_h$ .

Therefore, the following direct scheme for computing the numerical value of these functions for any set of initial conditions is proposed. Let  $E_2(\omega)$ , defined in Appendix 2, denote the lengthy expression for the terms of order  $\varepsilon^2$  in (2.8). Even though the form shown in the Appendix contains  $\phi$  explicitly, it was pointed out that upon simplification the terms depending upon  $\phi$  must cancel identically. Thus, strictly speaking,  $E_2$  is a function of  $\omega$  and the two initial values  $e_0$  and  $e_0$ . Furthermore, it is relatively straightforward to show that upon simplification  $E_2$  will reduce to the following form

$$\eta_{o}(e_{o}, i_{oo}) + \eta_{2}(e_{o}, i_{oo}) \cos 2\omega + \eta_{4}(e_{o}, i_{oo}) \cos 4\omega = E_{2}(\omega, e_{o}, i_{oo})$$
(2.10)

The next step consisting of exhibiting the explicit dependence of the  $\eta_i$  upon the initial values  $e_0$ , and  $i_{00}$ , is very arduous. The problem of predicting the numerical value of the  $\eta_i$  for any given pair  $e_0$ , and  $i_{00}$ , is however quite

simple. For any value of  $\omega$  and a given pair  $e_0$ , and  $i_{00}$ , (2.10) is a linear algebraic equation for the three  $\eta_i$ . The numerical value of the right hand side can be computed from the definition of  $E_2$  in Appendix 2 while the left hand side contains two coefficients which depend on  $\omega$ . Thus, if three distinct values of  $\omega$  are chosen, (2.10) produces three linear equations for the  $\eta_i$ .

One may, for instance, choose to evaluate  $E_2$  at  $\omega=0$ ,  $\frac{\pi}{4}$ , and  $\frac{\pi}{2}$ , to simplify the left-hand side of (2.10). Moreover, since  $E_2$  does not depend on  $\phi$  in its final form, a choice of  $\phi$  which results in considerable simplification of the numerical work of evaluating the right hand side of (2.10) is  $\phi=0$ .

Thus, a possible set of equations that results are

$$\eta_0 + \eta_2 + \eta_4 = E_2(0)$$
 (a)

$$\eta_{o} - \eta_{\downarrow} = E_{2}(\frac{\pi}{4})$$
 (2.11)

$$\eta_{0} - \eta_{2} + \eta_{4} = E_{2}(\frac{\pi}{2}) \tag{c}$$

The solution of the above system gives

$$n_{o} = \frac{1}{4} \left[ E_{2}(0) + 2E_{2}(\frac{\pi}{4}) + E_{2}(\frac{\pi}{2}) \right]$$
 (a)

It is allowable to choose arbitrary values for  $\omega$  since (2.10) is an algebraic identity even though the range of  $\omega$ , as predicted by the solution, is restricted for the case of critical inclination.

$$\eta_2 = \frac{1}{2} \left[ \mathbb{E}_2(0) - \mathbb{E}_2(\frac{\pi}{2}) \right]$$
 (2.12)

$$\eta_{\downarrow} = \frac{1}{4} \left[ E_2(0) - 2E_2(\frac{\pi}{4}) + E_2(\frac{\pi}{2}) \right] \tag{c}$$

Equation (2.9) can now be used to express the leading term of the Fourier series for  $\cos i_0 / pu_0^2$  in terms of known quantities in the form:

$$\frac{p^{3}}{\cos^{3}i_{o}(1-e^{2})^{3/2}} = (-2E)^{-3/2} + \frac{3}{2} (-2E)^{-5/2} \frac{\cos^{6}i_{o}}{p^{6}} [(\frac{1}{3}+e^{2})(1-3\cos^{2}i_{o}) + \frac{e^{2}}{4} (1+\cos^{2}i_{o})\cos 2\omega]$$

$$+ \epsilon^{2} \{+\frac{3}{2} (-2E)^{-5/2} [\eta_{o} + \eta_{2}\cos 2\omega + \eta_{4}\cos 4\omega] + \frac{3}{8} (-2E)^{-7/2} \frac{\cos^{12}i_{o}}{p^{12}} [(\frac{1}{3}+e^{2})^{2}(1-3\cos^{2}i_{o})^{2} + \frac{e^{2}}{2} (\frac{1}{3}+e^{2})(1-2\cos^{2}i_{o}-3\cos^{4}i_{o})\cos 2\omega$$

$$+ \frac{e^{4}}{22} (1+\cos^{2}i_{o})^{2}(1+\cos 4\omega)]\} \qquad (2.13)$$

# 2.3 The First Order Term

With the results of the previous section and use of the known solutions for  $u_0$ ,  $u_1$ , and  $i_1$ , the term of order  $\epsilon$  in the differential equation for the time becomes

$$T_{s}^{(1)} + T_{l}^{(1)} = \frac{\varepsilon \cos i_{o}}{p[1 + e \cos(\phi - \omega)]^{3}} \{1 + \frac{e^{2}}{2} - (4 + 3e^{2}) \cos^{2}i_{o} + (-\frac{1}{2} + \frac{5}{4} \cos^{2}i_{o})e^{2} \cos 2\omega - 2e \cos^{2}i_{o} \cos(\phi - \omega)$$

$$- \frac{e}{4} (3 - 7 \cos^{2}i_{o}) \cos(\phi + \omega) + [-\frac{1}{6} (1 - e^{2})$$

$$+ \frac{7}{6} \cos^{2}i_{o}] \cos 2\phi - \frac{e^{2}}{6} (1 - 6 \cos^{2}i_{o}) \cos(2\phi - 2\omega)$$

$$+ \frac{1}{2} e \cos^{2}i_{o} \cos(3\phi - \omega) + \frac{e^{2}}{12} \cos^{2}i_{o} \cos(4\phi - 2\omega)$$

$$+ \frac{3}{2} \varepsilon (-2E)^{-5/2} \frac{\cos^{6}i_{o}}{p^{6}} [(\frac{1}{3} + e^{2})(1 - 3 \cos^{2}i_{o})$$

$$+ \frac{e^{2}}{4} (1 + \cos^{2}i_{o}) \cos 2\omega] \qquad (2.14)$$

The notation  $T_s^{(1)}$  and  $T_l^{(1)}$  has been introduced to denote the fact that the right hand side of (2.14) consists of two parts. One is purely periodic in  $\phi$  and is denoted by  $T_s^{(1)}$  while the remainder, expressed as  $T_l^{(1)}$ , depends only on  $\tilde{\phi}$  (or  $\bar{\phi}$ ). The terms comprising  $T_l^{(1)}$  are functions of e, i<sub>0</sub>, and  $\omega$ , and will be studied separately because they give rise to either secular or long-period terms (dropping in order by one power of  $\varepsilon$ , after integration in the time solution). In order to effect the separation of the short-period terms, the Fourier expansion of  $[1 + e \cos(\phi - \omega)]^{-3}$ , given in Appendix 3, is used. It can then be shown that  $T_l^{(1)}$  becomes

$$T_{\ell}^{(1)}(e, i_{0}, \omega) = \frac{\cos i_{0}}{2p} \left\{ b_{30} \left[ 1 + \frac{e^{2}}{2} - (4 + 3e^{2}) \cos^{2} i_{0} \right] - \frac{e^{2}}{4} (2 - 5 \cos^{2} i_{0}) \cos 2\omega \right\}$$

$$- b_{31} \left[ 2e \cos^{2} i_{0} + \frac{e}{4} (3 - 7 \cos^{2} i_{0}) \cos 2\omega \right]$$

$$+ b_{32} \left[ -\frac{e^{2}}{6} (1 - 6 \cos^{2} i_{0}) - \frac{1}{6} (1 - e^{2} - 7 \cos^{2} i_{0}) \cos 2\omega \right]$$

$$+ \frac{1}{2} b_{33} e \cos^{2} i_{0} \cos 2\omega + \frac{1}{12} b_{34} e^{2} \cos^{2} i_{0} \cos 2\omega \right\}$$

$$+ \frac{3}{2} (-2E)^{-5/2} \frac{\cos^{6} i_{0}}{p^{6}} \left[ (\frac{1}{3} + e^{2})(1 - 3\cos^{2} i_{0}) + \frac{e^{2}}{4} (1 + \cos^{2} i_{0}) \cos 2\omega \right]$$

$$(2.15)$$

Note that since  $T_{\ell}^{(1)}$  comprises the long-period terms and is multiplied by  $\epsilon$  in (2.3) one must evaluate it correct to  $O(\epsilon)$  in order to insure the validity of the differential equation for the time correct to  $O(\epsilon^2)$ . This can be achieved for the case of critical inclination by substituting the known expansions for e and  $I_{0}$  (cf. (3.28b\*), (3.35\*), (3.37\*) and (3.28a\*), (3.33\*), (3.36\*)) giving the following formal representation of  $T_{\ell}^{(1)}$ 

$$T_{\ell}^{(1)}(e, i_{o}, \omega) = T_{\ell}^{(1)*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}, \varepsilon)$$

$$= T_{\ell o}^{(1)*}(e_{o}^{*}, i_{oo}^{*}, \varepsilon) + T_{\ell 1}^{(1)*}(e_{o}^{*}, i_{oo}^{*}, \varepsilon) \sqrt{\overline{\kappa_{o} - \kappa_{1}} \cos 2\omega_{c}}$$

$$+ T_{\ell 2}^{(1)*}(e_{o}^{*}, i_{oo}^{*}, \epsilon)(\sqrt{\kappa_{o} - \kappa_{1} \cos 2\omega_{c}})^{2}$$

$$+ T_{\ell 3}^{(1)*}(e_{o}^{*}, i_{oo}^{*}, \epsilon)(\sqrt{\kappa_{o} - \kappa_{1} \cos 2\omega_{c}})^{3}$$

$$+ T_{\ell h}^{(1)*}(e_{o}^{*}, i_{oo}^{*}, \epsilon)(\sqrt{\kappa_{o} - \kappa_{1} \cos 2\omega_{c}})^{4} \qquad (2.16)$$

Again, the explicit evaluation of the coefficients  $T_{\ell,i}^{(1)*}$ , i=0,1,...4 is a formidable task. However, a numerical procedure identical to the one proposed in section 2.2 can be used.

This consists of numerically evaluating  $T_{\ell}^{(1)}$  for any given pair of initial values  $e_0$  and  $i_{00}$  for five different and consistent (cf. what follows) values of  $\cos 2\omega_c$ , say  $\cos 2\omega_c^{(\nu)}$ ,  $\nu = 0, 1, ...4$ .

Clearly, the values of  $\omega_{\rm c}^{(\nu)}$  must be consistent with the permissible range given by the solution near the critical inclination, since now the explicit form of (2.19) depends on whether or not  $i_{\rm co}$  is near the critical value. It was shown in (3.51\*) and in(3.52\*) that near the critical inclination  $\omega_{\rm c}^{(\nu)}$  must obey the inequality

$$\sin^2 \omega_c^{(\nu)} > \frac{\kappa_1 - \overline{\kappa}_o}{2\kappa_1} \tag{2.17}$$

and possible choices of  $\omega_{\rm c}^{(\nu)}$  are

$$\omega_{c}^{(\nu)} = \frac{1}{2} \cos^{-1} \left[1 - \frac{\nu \cdot \sqrt{\overline{\kappa}_{o} - \kappa_{1}}}{\sqrt{\overline{\kappa}_{o} + \kappa_{1}} + \sqrt{\overline{\kappa}_{o} - \kappa_{1}}} - \frac{\nu^{2}}{4} \frac{\kappa_{1}}{(\sqrt{\overline{\kappa}_{o} + \kappa_{1}} + \sqrt{\overline{\kappa}_{o} - \kappa_{1}})^{2}}\right]$$
if  $\overline{\kappa}_{o} > \kappa_{1}$  (2.18)

or

$$\omega_{c}^{(\nu)} = \frac{1}{2} \cos^{-1} \left[ \frac{\overline{\kappa}_{o}}{\kappa_{1}} - \nu^{2} \left( \frac{\overline{\kappa}_{o}}{\kappa_{1}} + 1 \right) \right] \quad \text{if } -\kappa_{1} < \overline{\kappa}_{o} \le \kappa_{1}$$
 (2.19)

When each of the above values of  $\omega_c^{(v)}$  is substituted into the right hand side of (2.15) one obtains a value for the corresponding  $T_L^{(1)*}$  which can then be used in conjunction with (2.16) to derive a set of five linear algebraic equations for the five unknown  $T_L^{(1)*}$ . The results of such a calculation are summarized in Appendix 4.

If  $i_{00}$  is not critical the form of  $T_{\ell}^{(1)*}$  is simplified to

$$T_{\ell}^{(1)} (e, i_{o}, \omega) = T_{\ell}^{(1)} (e_{o}, i_{oo}, \omega, \varepsilon)$$

$$= T_{\ell o}^{(1)} (e_{o}, i_{oo}, \varepsilon) + T_{\ell 2}^{(1)} (e_{o}, i_{oo}, \varepsilon) \cos 2\omega$$

$$+ T_{\ell 4}^{(1)} (e_{o}, i_{oo}, \varepsilon) \cos 4\omega \quad (2.20.$$

In this case there are only three unknown functions:  $T_{\ell o}^{(1)}$ ,  $T_{\ell 2}^{(1)}$ , and  $T_{\ell h}^{(1)}$ . Since  $\omega$  undergoes a secular variation in this case one may use the following three equally spaced values of  $\omega=0$ ,  $\frac{\pi}{4}$ ,  $\frac{\pi}{2}$  to obtain the following system of three linear algebraic equations.

$$T_{\ell O}^{(1)} + T_{\ell 2}^{(1)} + T_{\ell 4}^{(1)} = T_{\ell}^{(1)} (e_{o}, i_{oo}, 0, \varepsilon)$$

$$T_{\ell O}^{(1)} - T_{\ell 4}^{(1)} = T_{\ell}^{(1)} (e_{o}, i_{oo}, \frac{\pi}{4}, \varepsilon)$$

$$T_{\ell O}^{(1)} - T_{\ell 2}^{(1)} + T_{\ell 4}^{(1)} = T_{\ell}^{(1)} (e_{o}, i_{oo}, \frac{\pi}{2}, \varepsilon)$$

$$(2.21)$$

The solution of (2.21) gives

$$T_{\ell o}^{(1)} = \frac{1}{l_{\ell}} [T_{\ell}^{(1)} (e_{o}, i_{oo}, 0, \varepsilon) + 2T_{\ell}^{(1)} (e_{o}, i_{oo}, \frac{\pi}{l_{\ell}}, \varepsilon) + T_{\ell}^{(1)} (e_{o}, i_{oo}, \frac{\pi}{2}, \varepsilon)]$$

$$T_{\ell 2}^{(1)} = \frac{1}{2} [T_{\ell}^{(1)} (e_{o}, i_{oo}, 0, \varepsilon) - T_{\ell}^{(1)} (e_{o}, i_{oo}, \frac{\pi}{2}, \varepsilon)]$$

$$T_{\ell l_{\ell}}^{(1)} = \frac{1}{l_{\ell}} [T_{\ell}^{(1)} (e_{o}, i_{oo}, 0, \varepsilon) - 2T_{\ell}^{(1)} (e_{o}, i_{oo}, \frac{\pi}{l_{\ell}}, \varepsilon) + T_{\ell}^{(1)} (e_{o}, i_{oo}, \frac{\pi}{2}, \varepsilon)]$$

$$(2.22)$$

With the definition given in Appendix 6 for the short-period terms to  $O(\epsilon)$ , the results are in a form suitable for quadrature. This will be discussed in Section 2.5 after the terms of  $O(\epsilon^2)$  have also been calculated.

# 2.4 The Second Order Term

For abbreviation, the whole expression to order  $\epsilon^2$  of (2.3) plus the terms of order  $\epsilon^2$  of the right side of (2.12), are denoted by  $T^{(2)}(i_{00}, e_0, \omega, \phi)$ ,

(cf. Appendix 5). When the solutions for  $u_0$ ,  $u_1$ ,  $u_2$ ,  $i_0$ ,  $i_1$ ,  $i_2$  are substituted and only the leading terms  $e_0$ ,  $i_0$  are kept, one may express  $T^{(2)}$  as a Fourier series in the form

$$T^{(2)}(i_{00}, e_{0}, \omega, \phi) = \frac{1}{2} T_{0}^{(2)}(i_{00}, e_{0}, \omega) + \sum_{m=1}^{\infty} T_{m}^{(2)}(i_{00}, e_{0}, \omega) \cos m(\phi - \omega)$$
(2.23)

Only the first term,  $\frac{1}{2}T^{(2)}(i_{00}, e_{0}, \omega)$ , is of interest because this is the only term that contributes to order  $\varepsilon$  in the solution for the time t. From inspection of the expression for  $T^{(2)}$ , it can be shown that  $T_{0}^{(2)}(i_{00}, e_{0}, \omega)$  is of the form

$$T_{o}^{(2)}(i_{oo}, e_{o}, \omega) = T_{oo}^{(2)}(i_{oo}, e_{o}) + T_{o2}^{(2)}(i_{oo}, e_{o}) \cos 2\omega + T_{o4}^{(2)}(i_{oo}, e_{o}) \cos 4\omega$$
(2.24)

The coefficients  $T_{00}^{(2)}$ ,  $T_{02}^{(2)}$ , and  $T_{04}^{(2)}$  can be evaluated by substitution of three different values for  $\omega$  in  $T_{0}^{(2)}(\omega)$ . The procedure is analogous to the one used for equations (2.10) and (2.11) and the solution is

$$T_{oo}^{(2)}(i_{oo}, e_{o}) = \frac{1}{4} [T_{o}^{(2)}(i_{oo}, e_{o}, 0) + 2 T_{o}^{(2)}(i_{oo}, e_{o}, \frac{\pi}{4}) + T_{o}^{(2)}(i_{oo}, e_{o}, \frac{\pi}{2})]$$

$$T_{o2}^{(2)}(i_{oo}, e_{o}) = \frac{1}{2} [T_{o}^{(2)}(i_{oo}, e_{o}, 0) - T_{o}^{(2)}(i_{oo}, e_{o}, \frac{\pi}{2})]$$

$$T_{0\downarrow}^{(2)}(i_{00}, e_{0}) = \frac{1}{1} \left[ T_{0}^{(2)}(i_{00}, e_{0}, 0) - 2 T_{0}^{(2)}(i_{00}, e_{0}, \frac{\pi}{4}) + T_{0}^{(2)}(i_{00}, e_{0}, \frac{\pi}{2}) \right]$$
(2.25)

where

$$T_o^{(2)}(i_{oo}, e_o, \omega) = \frac{1}{2\pi} \int_{-\pi}^{+\pi} T^{(2)}(i_{oo}, e_o, \omega, \phi) d\phi$$
 (2.26)

The definite integral of (2.26) has to be evaluated numerically for  $\omega = 0$ ,  $\frac{\pi}{4}$  and  $\frac{\pi}{2}$ , with the  $e_0$  and  $i_{00}$  given by the initial conditions.

# 2.5 Integration of the Time Equation

Using the results of the previous sections equation (2.3) may now be written in the following form:

$$\frac{dt}{d\phi} = (-2E)^{-3/2} + \frac{p^3}{\cos^3 i} \sum_{k=1}^{\infty} b_{2k} \cos k(\phi - \omega) 
+ \varepsilon \left[T_{lo}^{(1)*} + T_{ll}^{(1)*} \left[\overline{\kappa}_o - \kappa_l \cos 2\omega_c\right]^{1/2} + T_{l2}^{(1)*} \left(\overline{\kappa}_o - \kappa_l \cos 2\omega_c\right) 
+ T_{l3}^{(1)*} \left[\overline{\kappa}_o - \kappa_l \cos 2\omega_c\right]^{3/2} + T_{l4}^{(1)*} \left(\overline{\kappa}_o - \kappa_l \cos 2\omega_c\right)^2 + \frac{d\sigma}{d\phi} \right] 
+ \varepsilon^2 \left[T_{l0}^{(2)} + T_{l2}^{(2)} \cos 2\omega_c + T_{l4}^{(2)} \cos 4\omega_c\right] + O(\varepsilon^{5/2})$$
(2.27)

where  $\frac{d\sigma}{d\phi}$  stands for all the short period terms arising from the expression in (2.3) to order  $\varepsilon$ . This equation holds for all inclinations. For inclinations away from the critical, the following somewhat simpler form results

$$\frac{dt}{d\phi} = (-2E)^{-3/2} + \frac{p^3}{\cos^3 i_0} \sum_{k=1}^{\infty} b_{2k} \cos k(\phi - \omega) + \epsilon [T_{lo}^{(1)} + T_{l2}^{(1)} \cos 2\omega + T_{lk}^{(1)} \cos 4\omega] + \frac{d\sigma}{d\phi} + \epsilon^2 [T_{lo}^{(2)} + T_{lo}^{(2)} \cos 2\omega + T_{ok}^{(2)} \cos 4\omega] + 0(\epsilon^3)$$
(2.28)

and is applicable to most cases.

Integration of (2.28) yields

$$t = (-2E)^{-3/2}\phi + \frac{1}{2}\frac{T_{k2}^{(1)}}{S_o}\sin 2\omega + \frac{1}{4}\frac{T_{kk}^{(1)}}{S_o}\sin 4\omega + \frac{p^3}{\cos^3 i_o}\sum_{k=1}^{\infty}\frac{b_{2k}}{k}\sin k(\phi - \omega)$$

$$+ \varepsilon[T_{ko}^{(1)}\phi + \frac{1}{2}\frac{T_{o2}^{(2)}}{S_o}\sin 2\omega + \frac{1}{4}\frac{T_{ok}^{(2)}}{S_o}\sin 4\omega + \frac{S_op^3}{\cos^3 i_o}\sum_{k=1}^{\infty}\frac{b_{2k}}{k}\sin k(\phi - \omega)$$

$$+ \sigma(\phi, \omega)] + \varepsilon^2T_{oo}^{(2)} \cdot \phi + t_o$$
(2.29)

where  $\sigma$  is given in Appendix 6 and to is a constant to be determined by the initial conditions. Integration of the general equation yields the more complicated expression

$$t = (-2E)^{-3/2}\phi + \varepsilon^{-1/2} T_{\ell 1}^{(1)} (\omega_o^* - w) + \varepsilon^{-1/2} \int_{w}^{\omega_c} (\overline{\kappa}_o - \kappa_1 \cos 2\xi)^{1/2} [T_{\ell 2}^{(1)*} + T_{\ell 3}^{(1)*} \sqrt{\overline{\kappa}_o - \kappa_1 \cos 2\xi} + T_{\ell 4}^{(1)*} (\overline{\kappa}_o - \kappa_1 \cos 2\xi)] d\xi$$

$$+ \frac{p^{3}}{\cos^{3} i_{o}} \sum_{k=1}^{\infty} \frac{b_{2k}}{k} \sin k(\phi - \omega_{c}) + \varepsilon^{1/2} \int_{w}^{\omega} c \frac{T_{02}^{(2)} \cos 2\xi + T_{04}^{(2)} \cos 4\xi}{(\kappa_{o} - \kappa_{1} \cos 2\xi)^{1/2}} d\xi$$

$$+ \varepsilon [T_{k0}^{(1)} + \phi + \sigma] + \varepsilon^{2} T_{00}^{(2)} \cdot \phi + t_{0}^{*}$$
(2.30)

where to is the integration constant, to be determined by the initial conditions.

The extremely slow motion of the apse in case of critical inclination is responsible for the occurrence of the large terms of order  $\varepsilon^{-1/2}$  appearing in (2.30). However, it should be noted that these terms, although multiplied by  $\varepsilon^{-1/2}$ , are of order  $\varepsilon$  for  $\phi$  = 0(1) and grow to order  $\varepsilon^{-1/2}$  only when  $(\omega_{\rm c} - {\bf w}) = 0(1)$  which is equivalent to  $\phi$  = 0( $\varepsilon^{-3/2}$ ). The evaluation of the integrals occurring in (2.30) leads to elliptic functions and will not be exhibited here.

## REFERENCES

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- 2. Struble, R. A., <u>Journal of Mathematical Analysis and Application</u>, Vol. 1, No. 3 & 4, p. 300, 1960.
- 3. \_\_\_\_, Journal of Mathematics and Mechanics, Vol. 10, p. 691, 1962.

$$\begin{split} \mathbf{i}_{2} &= \frac{\cos^{9} i_{00} \sin i_{00}}{p^{8}} \left( -\frac{e_{0}}{48} \left( 19 - 21 \cos^{2} i_{00} \right) \cos(\phi - \omega) \right. \\ &- \frac{e_{0}}{4} \left[ - 1 + 9 \cos^{2} i_{00} + 6c \left( 1 + \frac{e_{0}^{2}}{4} \right) \left( 1 - 7 \cos^{2} i_{00} \right) \right] \cos(\phi + \omega) \\ &+ \frac{e_{0}^{3}}{8} c \left( 1 - 7 \cos^{2} i_{00} \right) \cos(\phi - 3\omega) + \frac{7}{16} e_{0}^{3} c \sin^{2} i_{00} \cos(\phi + 3\omega) \\ &- \frac{1}{4} \left[ 1 + \frac{e_{0}^{2}}{2} + 2c \left( 1 + \frac{3}{2} e_{0}^{2} \right) - \left( 1 + \frac{3}{2} e_{0}^{2} \right) \left( 1 + 14c \right) \cos^{2} i_{00} \right] \cos 2\phi \\ &- \frac{e_{0}^{2}}{24} \left( 2 - \cos^{2} i_{00} \right) \cos(2\phi - 2\omega) - \frac{e_{0}^{2}}{32} \left[ 5 - 7\cos^{2} i_{00} - 42c \sin^{2} i_{00} \right] \cos(2\phi + 2\omega) \\ &- \frac{e_{0}}{36} \left[ 1 + 7 \cos^{2} i_{00} + 18c \left( 1 + \frac{e_{0}^{2}}{4} \right) \left( 1 - 7 \cos^{2} i_{00} \right) \right] \cos(3\phi - \omega) \\ &- \frac{e_{0}}{24} \left[ 9 - 16 \cos^{2} i_{00} - 42c \left( 1 + \frac{e_{0}^{2}}{4} \right) \sin^{2} i_{00} \right] \cos\left( 3\phi + \omega \right) \\ &- \frac{1}{48} \left[ 10 \left( 1 + \frac{3}{4} e_{0}^{2} \right) - \left( 19 + \frac{27}{2} e_{0}^{2} \right) \cos^{2} i_{00} - 42c \left( 1 + \frac{3}{2} e_{0}^{2} \right) \sin^{2} i_{00} \right] \cos\left( 4\phi - 2\omega \right) \\ &+ \frac{e_{0}^{2}}{96} \left[ 1 - 15 \cos^{2} i_{00} - 18c \left( 1 - 7 \cos^{2} i_{00} \right) \right] \cos\left( 4\phi - 2\omega \right) \\ &- \frac{e_{0}}{48} \left[ 7 - 15 \cos^{2} i_{0} - \frac{252}{5} c \left( 1 + \frac{e_{0}^{2}}{4} \right) \sin^{2} i_{00} \right] \cos(5\phi - \omega) \\ &- \frac{e_{0}^{3}}{40} c \left( 1 - 7 \cos^{2} i_{0} \right) \cos(5\phi - 3\omega) \end{split}$$

$$-\frac{e_0^2}{288} \left[7 - 17 \cos^2 i_{00} - 126 c \sin^2 i_{00}\right] \cos(6\phi - 2\omega)$$

$$+\frac{e_0^3}{16} c \sin^2 i_{00} \cos(7\phi - 3\omega)\}$$

$$u_2 = \frac{\cos^{10} i_{00}}{p^{10}} \left\{-\frac{5}{24} - \frac{287}{288} e_0^2 + \frac{9}{8} c \left(1 + 3 e_0^2 + \frac{3}{8} e_0^{4}\right) + \left(-\frac{5}{8} + \frac{211}{48} e_0^2 + \frac{3}{8} e_0^2 + \frac{3}{8} e_0^2 + \frac{105}{8} e_0^2 + \frac$$

$$+ \left(\frac{245}{96} e_{o}^{2} - \frac{231}{8} c e_{o}^{2} - \frac{91}{16} c e_{o}^{4}\right) \cos^{4} i_{oo}] \cos \left(2\phi - 2\omega\right)$$

$$+ \left[-\frac{1}{168} e_{o}^{2} - \frac{7}{16} c e_{o}^{2} - \frac{7}{32} c e_{o}^{4} + \left(\frac{11}{148} e_{o}^{2} + \frac{7}{16} c e_{o}^{4}\right) \cos^{2} i_{oo} \right]$$

$$+ \left(-\frac{1}{6} e_{o}^{2} + \frac{7}{16} c e_{o}^{2} - \frac{7}{32} c e_{o}^{4}\right) \cos^{4} i_{oo}] \cos \left(2\phi + 2\omega\right)$$

$$+ \frac{1}{32} c e_{o}^{4} \left[1 - 24 \cos^{2} i_{o} + 35 \cos^{4} i_{oo}\right] \cos \left(2\phi + 4\omega\right)$$

$$+ \left[\frac{13}{288} e_{o} + \frac{9}{16} c e_{o} + \frac{e_{o}^{3}}{128} + \frac{31}{64} e_{o}^{3} c + \left(\frac{23}{24} e_{o} - \frac{13}{2} c e_{o} + \frac{25}{48} e_{o}^{3}\right)$$

$$- \frac{49}{8} c e_{o}^{3}\right) \cos^{2} i_{oo} + \left(-\frac{433}{288} e_{o} + \frac{119}{16} c e_{o} - \frac{85}{128} e_{o}^{3} \right)$$

$$+ \frac{469}{64} c e_{o}^{3}\right) \cos^{4} i_{oo}\right] \cos \left(3\phi - \omega\right) + \left[\frac{5}{96} e_{o} - \frac{7}{16} c e_{o} + \frac{25}{384} e_{o}^{3}\right]$$

$$- \frac{35}{64} c e_{o}^{3} + \left(-\frac{5}{48} e_{o} + \frac{7}{16} c e_{o} - \frac{51}{192} e_{o}^{3} + \frac{35}{32} c e_{o}^{3}\right) \cos^{2} i_{oo}$$

$$+ \left(+\frac{e_{o}}{12} + \frac{27}{128} e_{o}^{3} - \frac{35}{64} c e_{o}^{3}\right) \cos^{4} i_{oo}\right] \cos \left(3\phi + \omega\right)$$

+ 
$$\left[\frac{11}{384}e_{0}^{3} - \frac{9}{64}ce_{0}^{3} + \left(-\frac{73}{192}e_{0}^{3} + \frac{99}{32}ce_{0}^{3}\right)\cos^{2}i_{00}\right]$$

$$+ (\frac{125}{128} e_0^3 - \frac{357}{64} c e_0^3) cos^4 i_{00}] cos (3\phi - 3\omega) + [\frac{1}{24} - \frac{7}{40} c + \frac{13}{96} e_0^2]$$

$$-\frac{7}{10}ce_{o}^{2}-\frac{7}{64}ce_{o}^{4}+\left(-\frac{1}{6}+\frac{7}{20}c-\frac{7}{16}e_{o}^{2}+\frac{7}{4}ce_{o}^{2}-\frac{49}{160}ce_{o}^{4}\right)cos^{2}i_{oo}$$

+ 
$$(\frac{1}{8} - \frac{7}{40}c + \frac{23}{32}e_0^2 - \frac{21}{20}c e_0^2 - \frac{63}{320}c e_0^4) \cos^4 i_{00}] \cos 4\phi$$

$$+ \left[ -\frac{e^{2}}{36}^{2} + \frac{11}{10} c e^{2} + \frac{1}{20} c e^{4} + \left( +\frac{19}{24} e^{2} - \frac{21}{5} c e^{2} - \frac{1}{5} c e^{4} \right) \cos^{2} i_{00} \right]$$

$$+ \left( -\frac{11}{8} e^{2} + \frac{217}{40} c e^{2} + \frac{21}{20} c e^{4} \right) \cos^{4} i_{00} \right] \cos^{4} i_{00} \cos^{4} i_{00$$

$$\begin{split} E_{2}(\phi, \ \omega) &= \frac{p^{2}}{\cos^{2}i_{oo}} \left\{ - \left( \frac{\partial u_{1}}{\partial \phi} \right)^{2} - 2 \frac{\partial u_{1}}{\partial \phi} \frac{\partial u_{0}}{\partial \phi} - \left( \frac{\partial u_{0}}{\partial \phi} \right)^{2} - 2 \frac{\partial u_{2}}{\partial \phi} \frac{\partial u_{0}}{\partial \phi} - 2 \frac{\partial u_{1}}{\partial \phi} \frac{\partial u_{0}}{\partial \phi} \right. \\ &- 2 \frac{\sin 2i_{oo}}{p^{2}} \left( E + u_{0} \right) i_{2} - 2 \frac{\cos 2i_{oo}}{p^{2}} \left( E + u_{0} \right) i_{1}^{2} - 2 \frac{\sin 2i_{oo}}{p^{2}} u_{1} i_{1} \\ &+ 2 \left( \frac{\cos^{2}i_{oo}}{p^{2}} - u_{0} \right) u_{2} - u_{1}^{2} + u_{8} \frac{\cos^{5}i_{oo}}{p^{4}} \sin i_{oo} i_{1}u_{o} \left( E + u_{o} \right) \sin^{2}\phi \\ &- 8 \frac{\cos^{6}i_{oo}}{p^{2}} \left( E + 2u_{o} \right) u_{1} \sin^{2}\phi - 16 \frac{\cos^{3}i_{oo}}{p^{2}} \sin i_{oo} i_{1}u_{o}^{3} \sin^{2}\phi \\ &+ 12 \frac{\cos^{4}i_{oo}}{p^{2}} u_{o}^{2} u_{1} \sin^{2}\phi - \frac{2}{3} \frac{\sin 2i_{oo}}{p^{2}} i_{1}u_{o}^{3} - \frac{\sin^{4}i_{o}}{p^{2}} i_{1}u_{o}^{3} \sin^{2}\phi \\ &+ 2 \frac{\cos^{2}i_{oo}}{p^{2}} u_{o}^{2} u_{1} (1 - 3 \sin^{2}i_{oo} \sin^{2}\phi) \\ &+ 12 \frac{\cos^{12}i_{oo}}{p^{8}} e_{o}^{2} u_{o}^{2} \sin^{2}(\phi - \omega) \sin^{4}\phi \\ &+ 56 c \frac{\cos^{8}i_{oo}}{p^{6}} e_{o}^{2} u_{o}^{3} \sin^{2}(\phi - \omega) \sin^{4}\phi \\ &- 2^{4} c \frac{\cos^{8}i_{oo}}{p^{6}} e_{o}^{2} u_{o}^{3} \sin^{2}(\phi - \omega) \sin^{2}\phi \\ &- \frac{3}{3} \frac{\cos^{6}i_{oo}}{p^{6}} u_{o}^{4} \left( \sin^{2}\phi - 3 \sin^{2}i_{oo} \sin^{4}\phi \right) \\ &+ 14 c \frac{\cos^{2}i_{oo}}{p^{2}} u_{o}^{5} \sin^{4}i_{oo} \sin^{4}\phi - 12 c \frac{\cos^{2}i_{oo}}{p^{2}} u_{o}^{5} \sin^{2}i_{oo} \sin^{2}\phi \\ &+ \frac{6}{5} c \frac{\cos^{2}i_{oo}}{2} u_{o}^{5} \sin^{4}i_{oo} \sin^{4}\phi - 12 c \frac{\cos^{2}i_{oo}}{p^{2}} u_{o}^{5} \sin^{2}i_{oo} \sin^{2}\phi \\ &+ \frac{6}{5} c \frac{\cos^{2}i_{oo}}{2} u_{o}^{5} \end{array}$$

$$-\frac{\cos^{6}i_{00}}{p^{8}}i_{1}\sin^{2}2i_{00}[(\frac{1}{3}+e_{0}^{2})(1-3\cos^{2}i_{00})$$

$$+\frac{e^{2}}{4}(1+\cos^{2}i_{00})\cos^{2}2\omega]$$

$$-4\frac{\cos^{12}i_{00}}{p^{10}}u_{0}\sin^{2}\phi[(\frac{1}{3}+e_{0}^{2})(1-3\cos^{2}i_{00})$$

$$+\frac{e^{2}}{4}(1+\cos^{2}i_{00})\cos^{2}2\omega]$$

$$+2\frac{\partial u_{0}}{\partial \phi}\frac{\sin^{2}2i_{00}}{p^{2}}c_{2}\sin^{2}2\omega[1+e_{0}\cos(\phi-\omega)]$$

$$-2\frac{\partial u_{0}}{\partial \phi}\frac{\cos^{2}i_{00}}{p^{2}}B_{2}\sin^{2}2\omega\cos(\phi-\omega)$$

Fourier Expansion of the Function

$$\frac{1}{\left[1+e\cos(\phi-\omega)\right]^n}$$

$$\frac{1}{\left[1+e\cos(\phi-\omega)\right]^n}=\frac{b_{no}}{2}+\sum_{k=1}^{\infty}b_{nk}\cos k(\phi-\omega)$$

where, for k = 0, 1, 2 ...

$$b_{1k} = \frac{2}{\sqrt{1 - e^2}} \left( \frac{\sqrt{1 - e^2} - 1}{e} \right)^k$$

$$b_{2k} = \frac{2}{(1 - e^2)^{3/2}} \left( \frac{\sqrt{1 - e^2} - 1}{e} \right)^k (1 + k \sqrt{1 - e^2})$$

$$b_{3k} = \frac{2}{(1 - e^2)^{5/2}} \left( \frac{\sqrt{1 - e^2} - 1}{e} \right)^k \left( 1 + \frac{e^2}{2} + \frac{3}{2} k \sqrt{1 - e^2} + \frac{k^2}{2} (1 - e^2) \right)$$

$$b_{4k} = \frac{2}{(1 - e^2)^{7/2}} \left( \frac{\sqrt{1 - e^2 - 1}}{e} \right)^k \left[ 1 + \frac{3}{2} e^2 + \frac{11 + 4}{6} e^2 \right] k \sqrt{1 - e^2}$$

$$+ k^2 (1 - e^2) + \frac{k^3}{6} (1 - e^2)^{3/2}$$

$$T_{10}^{*}(e_{o}^{*},i_{oo}^{*},\epsilon) = [(1+\frac{25}{12}y+\frac{35}{24}y^{2}+\frac{5}{12}y^{3}+\frac{1}{24}y^{4}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(o)},\epsilon)$$

$$-(4y+\frac{13}{3}y^{2}+\frac{3}{2}y^{3}+\frac{1}{6}y^{4}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(1)},\epsilon)$$

$$+(3y+\frac{19}{4}y^{2}+2y^{3}+\frac{1}{4}y^{4}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(2)},\epsilon)$$

$$-(\frac{1}{3}y+\frac{7}{3}y^{2}+\frac{7}{6}y^{3}+\frac{1}{6}y^{4}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(3)},\epsilon)$$

$$+(\frac{1}{4}y+\frac{11}{24}y^{2}+\frac{1}{4}y^{3}+\frac{1}{24}y^{4}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(4)},\epsilon)]$$

$$T_{11}^{*}(e_{o}^{*},i_{oo}^{*},\epsilon) = x[-(\frac{25}{12}+\frac{35}{12}y+\frac{5}{4}y^{2}+\frac{1}{6}y^{3}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(0)},\epsilon)$$

$$+(4+\frac{26}{3}y+\frac{9}{2}y^{2}+\frac{2}{3}y^{3}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(0)},\epsilon)$$

$$-(3+\frac{19}{2}y+6y^{2}+y^{3}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(2)},\epsilon)$$

$$+(\frac{1}{3}+\frac{11}{3}y+\frac{7}{2}y^{2}+\frac{2}{3}y^{3}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(3)},\epsilon)$$

$$-(\frac{1}{1}+\frac{11}{12}y+\frac{3}{1}y^{2}+\frac{1}{6}y^{3}) T_{1}^{*}(e_{o}^{*},i_{oo}^{*},\omega_{c}^{(3)},\epsilon)$$

$$T_{12}^{*}(e_{o}^{*}, i_{oo}^{*}, \varepsilon) = x^{2}[(\frac{35}{24} + \frac{5}{4}y + \frac{1}{4}y^{2}) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(o)}, \varepsilon)$$

$$- (\frac{13}{3} + \frac{9}{2}y + y^{2}) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(1)}, \varepsilon)$$

$$+ (\frac{19}{4} + 6y + \frac{3}{2}y^{2}) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(2)}, \varepsilon)$$

$$- (\frac{7}{3} + \frac{7}{2}y + y^{2}) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(3)}, \varepsilon)$$

$$+ (\frac{11}{24} + \frac{3}{4}y + \frac{1}{4}y^{2}) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(4)}, \varepsilon)]$$

$$T_{13}^{*}(e_{o}^{*}, i_{oo}^{*}, \varepsilon) = x^{3}[-(\frac{5}{12} + \frac{1}{6}y) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(o)}, \varepsilon) + (\frac{3}{2} + \frac{2}{3}y) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(1)}, \varepsilon)$$

$$-(2 + y) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(2)}, \varepsilon) + (\frac{7}{6} + \frac{2}{3}y) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(3)}, \varepsilon)$$

$$-(\frac{1}{4} + \frac{1}{6}y) T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(4)}, \varepsilon)]$$

$$T_{11}^{*}(e_{o}^{*}, i_{oo}^{*}, \epsilon) = x^{1/2} \left[ \frac{1}{21} T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(o)}, \epsilon) - \frac{1}{6} T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(1)}, \epsilon) \right]$$

$$+ \frac{1}{14} T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(2)}, \epsilon) - \frac{1}{6} T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(3)}, \epsilon)$$

$$+ \frac{1}{24} T_{1}^{*}(e_{o}^{*}, i_{oo}^{*}, \omega_{c}^{(4)}, \epsilon) \right]$$

where

$$x = \begin{cases} \frac{\frac{1}{\sqrt{\kappa_0 - \kappa_1}}, & \text{if } -\kappa_1 < \overline{\kappa_0} < +\kappa_1 \\ \\ \frac{1}{\sqrt{\kappa_0 + \kappa_1} - \sqrt{\kappa_0 - \kappa_1}}, & \text{if } \overline{\kappa_0} \ge \kappa_1 \end{cases}$$

$$y = \begin{cases} 0, & \text{if } -\kappa_1 < \overline{\kappa}_0 < +\kappa_1 \end{cases}$$

$$\frac{\sqrt{\overline{\kappa}_0 - \kappa_1}}{\sqrt{\overline{\kappa}_0 + \kappa_1} - \sqrt{\overline{\kappa}_0 - \kappa_1}}, & \text{if } \overline{\kappa}_0 \ge \kappa_1$$

$$T_{2}(i_{00}, e_{0}, \omega, \phi) = -\frac{1}{2} \frac{\cos i_{00}}{pu_{0}^{2}} i_{1}^{2} - \frac{\sin i_{00}}{pu_{0}^{2}} i_{2} + 2 \frac{\sin i_{00}}{pu_{0}^{3}} i_{1}^{u} 1$$

$$+ 3 \frac{\cos i_{00}}{pu_{0}^{4}} u_{1}^{2} - 2 \frac{\cos i_{00}}{pu_{0}^{3}} u_{2} + 10 \frac{\cos^{\frac{1}{3}} i_{0} \sin i_{00}}{p^{3} u_{0}^{3}} i_{1} \sin^{2} \phi$$

$$+ 2 \frac{\cos^{\frac{5}{3}} i_{00}}{p^{3} u_{0}^{2}} u_{1} \sin^{2} \phi + \frac{\cos^{\frac{9}{3}} i_{00}}{p^{5}} \sin^{\frac{1}{4}} \phi$$

$$+ 28 \frac{cu_{0}}{p^{3}} \cos^{\frac{5}{3}} i_{00} \sin^{2} i_{00} \sin^{\frac{1}{4}} \phi$$

$$- 12 \frac{cu_{0}}{p^{3}} \cos^{\frac{5}{3}} i_{00} \sin^{2} \phi$$

$$+ \frac{3}{2} (-2E)^{-\frac{5}{2}} [\eta_{0} + \eta_{2} \cos 2\omega + \eta_{1} \cos 4\omega]$$

$$+ \frac{3}{8} (-2E)^{-\frac{7}{2}} \frac{\cos^{\frac{12}{3}} i_{00}}{p^{\frac{1}{2}}} [(\frac{1}{3} + e_{0})^{2} (1 - 3 \cos^{2} i_{00})^{2}$$

$$+ \frac{e_{0}^{2}}{2} (\frac{1}{3} + e_{0}^{2}) (1 - 2 \cos^{2} i_{00} - 3 \cos^{\frac{1}{3}} i_{00}) \cos 2\omega$$

$$+ \frac{e_{0}^{4}}{32} (1 + \cos^{2} i_{00})^{2} (1 + \cos^{4} \omega)]$$

$$\begin{split} \sigma &= \frac{\cos i_0}{p} \sum_{k=1}^{\infty} b_{3k} \left[ \frac{1}{k} (1 + \frac{e^2}{2} - (4 + 3e^2) \cos^2 i_0) \sin k(\phi - \omega) \right. \\ &- \frac{e^2}{6k} (2 - 5 \cos^2 i_0) \sin(k\phi - (k - 2)\omega) - \frac{e}{k+1} \cos^2 i_0 \sin((k+1)(\phi - \omega)) \right. \\ &- \frac{e^2}{6k} (2 - 5 \cos^2 i_0) \sin(k\phi - (k+2)\omega) - \frac{e}{8} \frac{(3 - 7 \cos^2 i_0)}{k+1} \sin((k+1)\phi - (k-1)\omega) \\ &- \frac{1}{12(k+2)} (1 - e^2 - 7 \cos^2 i_0) \sin((k+2)\phi - k\omega) \\ &- \frac{e^2}{12} \frac{(1 - 6 \cos^2 i_0)}{(k+2)} \sin((k+2)(\phi - \omega)) \\ &+ \frac{e \cos^2 i_0}{4 + k + 3} \sin((k+3)\phi - (k+1)\omega) + \frac{e^2}{24} \frac{\cos^2 i_0}{(k+4)} \sin((k+4)\phi - (k+2)\omega) \right] \\ &- \frac{\cos i_0}{p} \sum_{k=2}^{\infty} b_{3k} \left[ \frac{e \cos^2 i_0}{k-1} \sin((k-1)\phi - \omega) + \frac{e}{8} \frac{(3 - 7 \cos^2 i_0)}{(k-1)} \sin((k-1)\phi - (k+1)\omega) \right] \\ &- \frac{\cos i_0}{p} \sum_{k=1}^{\infty} \frac{1}{12} \frac{b_{3k}}{(k-2)} \left[ - (1 - e^2 - 7 \cos^2 i_0) \sin((k-2)\phi - k\omega) + \frac{e^2}{p} (1 - 6 \cos^2 i_0) \sin((k-2)(\phi - \omega)) \right] \\ &+ \frac{\cos i_0}{p} \sum_{k=1}^{\infty} \frac{1}{4} \frac{b_{3k}}{(k-3)} e \cos^2 i_0 \sin((k-3)\phi - (k-1)\omega) \\ &+ \frac{\cos i_0}{p} \sum_{k=1}^{\infty} \frac{1}{24} \frac{b_{3k}}{(k-4)} e^2 \cos^2 i_0 \sin((k-3)\phi - (k-2)\omega) \end{aligned}$$

